Stability, Multiplicity, and Sun Spots (Blanchard-Kahn conditions)

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Introduction

- What do we mean with non-unique solutions?
 - multiple solution versus multiple steady states
- What are sun spots?
- Are models with sun spots scientific?

Terminology

- Definitions are very clear
 - (use in practice can be sloppy)

Model:

$$H(p_{+1},p)=0$$

Solution:

$$p_{+1} = f(p)$$

Unique solution & multiple steady states



Multiplicity

Examples

Multiple solutions & unique (non-zero) steady state



Multiple steady states & sometimes multiple solutions



From Den Haan (2007)

Examples

Large sun spots (around 2000 at the peak)



Sun spot cycle (almost at peak again)



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NOAA/SWPC Boulder,CO USA

Sun spots in economics

- **Definition:** a solution is a sun spot solution if it depends on a stochastic variable *from outside system*
- Model:

$$0 = \mathsf{E}H(p_{t+1}, p_t, d_{t+1}, d_t)$$

 d_t : exogenous random variable

Sun spots in economics

• Non-sun-spot solution:

$$p_t = f(p_{t-1}, p_{t-2}, \cdots, d_t, d_{t-1}, \cdots)$$

• Sun spot:

$$p_t = f(p_{t-1}, p_{t-2}, \cdots, d_t, d_{t-1}, \cdots, s_t)$$

$$s_t : \text{ random variable with } \mathsf{E}[s_{t+1}] = 0$$

Sun spots and science

Why are sun spots attractive

- sun spots: s_t matters, just because agents believe this
 - self-fulfilling expectations don't seem that unreasonable
- sun spots provide many sources of shocks
 - number of sizable fundamental shocks small

Sun spots and science

Why are sun spots not so attractive

- Purpose of science is to come up with predictions
 - If there is one sun spot solution, there are zillion others as well
- Support for the conditions that make them happen not overwhelming
 - you need sufficiently large increasing returns to scale or externality

Overview

Getting started

• simple examples

2 General derivation of Blanchard-Kahn solution

- When unique solution?
- When multiple solution?
- When no (stable) solution?
- 3 When do sun spots occur?
- O Numerical algorithms and sun spots

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General Derivation

Examples

Getting started

Model: $y_t = \rho y_{t-1}$

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General Derivation

Examples

Getting started

Model: $y_t = \rho y_{t-1}$

- infinite number of solutions, independent of the value of ho

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General Derivation

Examples

Getting started

Model: $\begin{array}{c} y_{t+1} = \rho y_t \\ y_0 \text{ is given} \end{array}$

General Derivation

Examples

Getting started

Model:
$$y_{t+1} = \rho y_t$$

 y_0 is given

- unique solution, independent of the value of ho

Getting started

• Blanchard-Kahn conditions apply to models that add as a requirement that the series do not explode

 $y_{t+1} =
ho y_t$ Model: y_t cannot explode

- $\rho > 1$: nique solution, namely $y_t = 0$ for all t
- $\rho < 1$: many solutions
- $\rho = 1$: many solutions
 - be careful with ho=1, uncertainty matters

State-space representation

$$Ay_{t+1} + By_t = \varepsilon_{t+1}$$
$$\mathsf{E} \left[\varepsilon_{t+1} | I_t \right] = 0$$
is an $n \times 1$ vector $y_t: m < n$ elements are not determined

some elements of ε_{t+1} are not exogenous shocks but prediction errors

Neoclassical growth model and state space representation

$$\mathsf{E} \begin{bmatrix} \left(\exp(z_t)k_{t-1}^{\alpha} + (1-\delta)k_{t-1} - k_t \right)^{-\gamma} = \\ \beta\left(\exp(z_{t+1})k_t^{\alpha} + (1-\delta)k_t - k_{t+1} \right)^{-\gamma} \\ \times \left(\alpha \exp\left(z_{t+1}\right)k_t^{\alpha-1} + 1 - \delta \right) \end{bmatrix} I_t \end{bmatrix}$$

or equivalently without $\mathsf{E}[\cdot]$

$$(\exp(z_t)k_{t-1}^{\alpha} + (1-\delta)k_{t-1} - k_t)^{-\gamma} = \beta (\exp(z_{t+1})k_t^{\alpha} + (1-\delta)k_t - k_{t+1})^{-\gamma} \times (\alpha \exp(z_{t+1})k_t^{\alpha-1} + 1 - \delta) + e_{\mathsf{E},t+1}$$

Neoclassical growth model and state space representation

Linearized model:

$$k_{t+1} = a_1k_t + a_2k_{t-1} + a_3z_{t+1} + a_4z_t + e_{\mathsf{E},t+1}$$

 $z_{t+1} = \rho z_t + e_{z,t+1}$
 k_0 is given

- k_t is end-of-period t capital
 - \implies k_t is chosen in t

Neoclassical growth model and state space representation

$$\begin{bmatrix} 1 & 0 & a_3 \\ 0 & 1 & 0 \\ 0 & 0 & 1 \end{bmatrix} \begin{bmatrix} k_{t+1} \\ k_t \\ z_{t+1} \end{bmatrix} + \begin{bmatrix} a_1 & a_2 & a_4 \\ -1 & 0 & 0 \\ 0 & 0 & -\rho \end{bmatrix} \begin{bmatrix} k_t \\ k_{t-1} \\ z_t \end{bmatrix} = \begin{bmatrix} e_{\mathsf{E},t+1} \\ 0 \\ e_{z,t+1} \end{bmatrix}$$

Examples

Dynamics of the state-space system

$$Ay_{t+1} + By_t = \varepsilon_{t+1}$$

$$y_{t+1} = -A^{-1}By_t + A^{-1}\varepsilon_{t+1}$$
$$= Dy_t + A^{-1}\varepsilon_{t+1}$$

Thus

$$y_{t+1} = D^t y_1 + \sum_{l=1}^{t+1} D^{t+1-l} A^{-1} \varepsilon_l$$

Jordan matrix decomposition

$$D = P\Lambda P^{-1}$$

- Λ is a diagonal matrix with the eigen values of D
- without loss of generality assume that $|\lambda_1| \ge |\lambda_2| \ge \cdots |\lambda_n|$

Let

$$P^{-1} = \begin{bmatrix} \tilde{p}_1 \\ \vdots \\ \tilde{p}_n \end{bmatrix}$$

where \tilde{p}_i is a $(1 \times n)$ vector

Examples

Dynamics of the state-space system

$$y_{t+1} = D^{t}y_{1} + \sum_{l=1}^{t+1} D^{t+1-l}A^{-1}\varepsilon_{l}$$
$$= P\Lambda^{t}P^{-1}y_{1} + \sum_{l=1}^{t+1} P\Lambda^{t+1-l}P^{-1}A^{-1}\varepsilon_{l}$$

Dynamics of the state-space system

multiplying dynamic state-space system with P^{-1} gives

$$P^{-1}y_{t+1} = \Lambda^t P^{-1}y_1 + \sum_{l=1}^{t+1} \Lambda^{t+1-l} P^{-1} A^{-1} \varepsilon_l$$

or

$$\tilde{p}_i y_{t+1} = \lambda_i^t \tilde{p}_i y_1 + \sum_{l=1}^{t+1} \lambda_i^{t+1-l} \tilde{p}_i A^{-1} \varepsilon_l$$

recall that y_t is $n \times 1$ and \tilde{p}_i is $1 \times n$. Thus, $\tilde{p}_i y_t$ is a scalar

Model

0
$$\tilde{p}_i y_{t+1} = \lambda_i^t \tilde{p}_i y_1 + \sum_{l=1}^{t+1} \lambda_i^{t+1-l} \tilde{p}_i A^{-1} \varepsilon_l$$

9 $\mathsf{E}[\varepsilon_{t+1}|I_t] = 0$

- **3** m elements of y_1 are not determined
- **4** y_t cannot explode

Reasons for multiplicity

- **1** There are free elements in y_1
- **2** The only constraint on $e_{E,t+1}$ is that it is a prediction error.
 - This leaves lots of freedom

Eigen values and multiplicity

- Suppose that $|\lambda_1| > 1$
- To avoid explosive behavior it *must* be the case that
- $\label{eq:planet} \begin{array}{ll} \mathbf{\hat{p}}_1 y_1 = 0 \quad \text{and} \\ \mathbf{\hat{p}}_1 A^{-1} \varepsilon_l = 0 \quad \forall l \end{array}$

How to think about #1?

$$\tilde{p}_1 y_1 = 0$$

- Simply an additional equation to pin down some of the free elements
- Much better: This is the policy rule in the first period

How to think about #1?

$$\tilde{p}_1 y_1 = 0$$

Neoclassical growth model:

- $y_1 = [k_1, k_0, z_1]^T$
- $|\lambda_1|>$ 1, $|\lambda_2|<$ 1, $\lambda_3=
 ho<$ 1
- $\tilde{p}_1 y_1$ pins down k_1 as a function of k_0 and z_1
 - this is the policy function in the first period

How to think about #2?

$$ilde{p}_1 A^{-1} arepsilon_l = 0 \;\; orall l$$

- This pins down $e_{E,t}$ as a function of $\varepsilon_{z,t}$
- That is, the prediction error must be a function of the structural shock, ε_{z,t}, and cannot be a function of other shocks,
 - i.e., there are no sunspots

How to think about #2?

$$ilde{
u}_1 A^{-1} arepsilon_l = 0 \;\; orall l$$

Neoclassical growth model:

• $\tilde{p}_1 A^{-1} \varepsilon_t$ says that the prediction error $e_{\mathsf{E},t}$ of period t is a fixed function of the innovation in period t of the exogenous process, $e_{z,t}$

Examples

How to think about #1 combined with #2?

$$\tilde{p}_1 y_t = 0 \quad \forall t$$

- Without sun spots
 - i.e. with $\tilde{p}_1 A^{-1} \varepsilon_t = 0 \;\; orall t$
- k_t is pinned down by k_{t-1} and z_t in every period.

Blanchard-Kahn conditions

- Uniqueness: For every free element in y_1 , you need one $\lambda_i > 1$
- Multiplicity: Not enough eigenvalues larger than one
- No stable solution: Too many eigenvalues larger than one

How come this is so simple?

• In practice, it is easy to get

$$Ay_{t+1} + By_t = \varepsilon_{t+1}$$

• How about the next step?

$$y_{t+1} = -A^{-1}By_t + A^{-1}\varepsilon_{t+1}$$

- **Bad news**: A is often not invertible
- Good news: Same set of results can be derived
 - Schur decomposition (See Klein 2000)

Examples

How to check in Dynare

Use the following command after the model & initial conditions part

check;

Examples

Example - x predetermined - 1st order

$$\begin{aligned} x_{t-1} &= E_t \left[\phi x_t + z_{t+1} \right] \\ z_t &= 0.9 z_{t-1} + \varepsilon_t \end{aligned}$$

- $|\phi| > 1$: Unique stable fixed point
- + $|\phi| < 1$: No stable solutions; too many eigenvalues > 1

Example - x predetermined - 2nd order

$$\phi_2 x_{t-1} = E_t [\phi_1 x_t + x_{t+1} + z_{t+1}] z_t = 0.9 z_{t-1} + \varepsilon_t$$

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$$\phi_1=-2.25,\,\phi_2=-0.5$$
 : Unique stable fixed point $(1+\phi_1L-\phi_2L^2)x=(1-2L)(1-\frac{1}{4}L)$

- $\phi_1 = -3.5$, $\phi_2 = -3$: No stable solution; too many eigenvalues > 1 $(1 + \phi_1 L - \phi_2 L^2)x = (1 - 2L)(1 - 1.5L)$
- $\phi_1 = -1$, $\phi_2 = -0.25$: Multiple stable solutions; too few eigenvalues > 1 $(1 + \phi_1 L - \phi_2 L^2)x = (1 - 0.5L)(1 - 0.5L)$

Examples

Example - x not predetermined - 1st order

$$x_t = E_t [\phi x_{t+1} + z_{t+1}]$$

$$z_t = 0.9z_{t-1} + \varepsilon_t$$

- $|\phi| < 1$: Unique stable fixed point
- + $|\phi|>1$: Multiple stable solutions; too few eigenvalues >1

References

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- Den Haan, Wouter J., 2007, Shocks and the Unavoidable Road to Higher Taxes and Higher Unemployment, Review of Economic Dynamics, 348-366.
 - simple model in which the size of the shocks has long-term consequences
- Farmer, Roger, 1993, The Macroeconomics of Self-Fulfilling Prophecies, The MIT Press.
 - textbook by the pioneer
- Klein, Paul, 2000, Using the Generalized Schur form to Solve a Multivariate Linear Rational Expectations Model
 - in case you want to do the analysis without the simplifying assumption that ${\cal A}$ is invertible